



BlueSi



25.04.2024



Martin Wendt (UP)



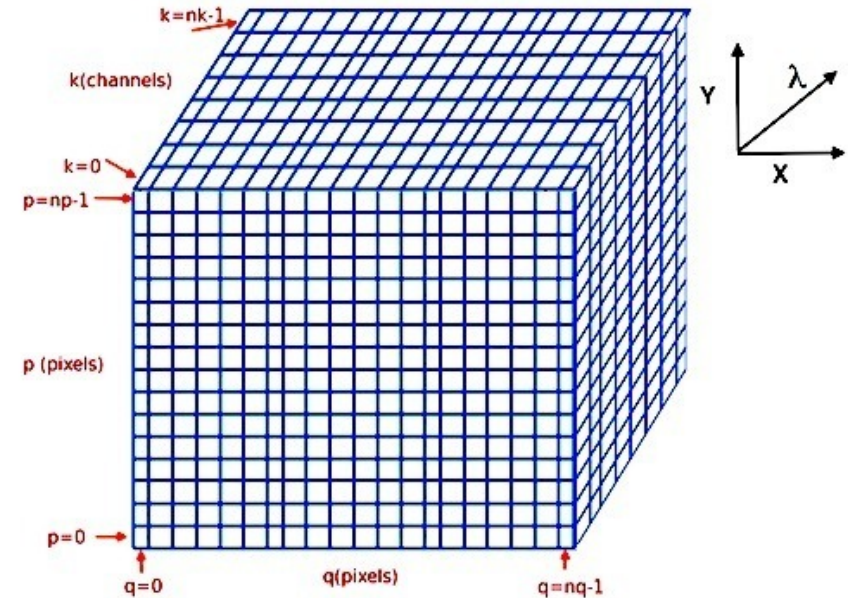
BlueSi:

Creating full scale and full feature cubes as expected from the BlueMUSE data reduction pipeline.

Technical specs as close as feasible.
All data content is simulated.

Level of detail on cube level (*)
(no CCD simulation or light path rendering)

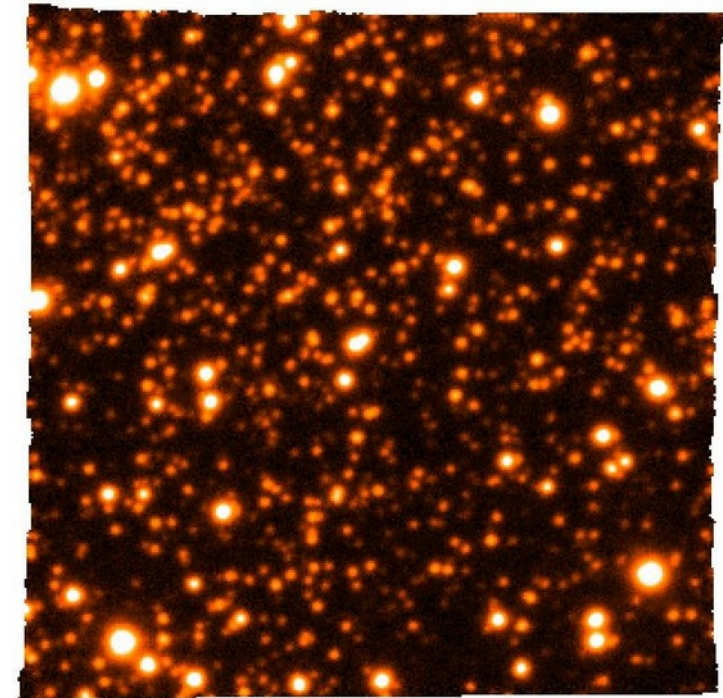
BUT including environmental conditions.

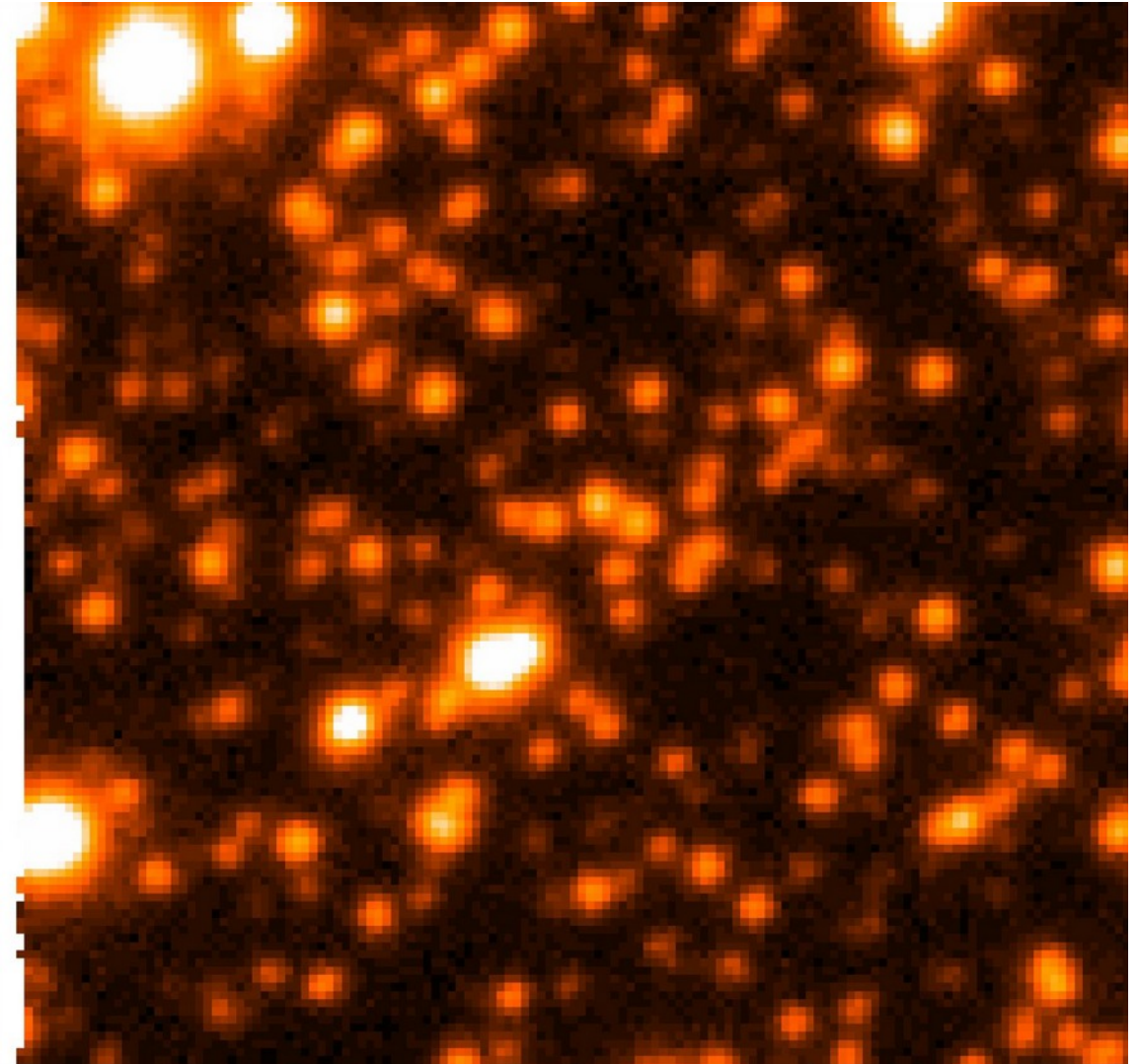
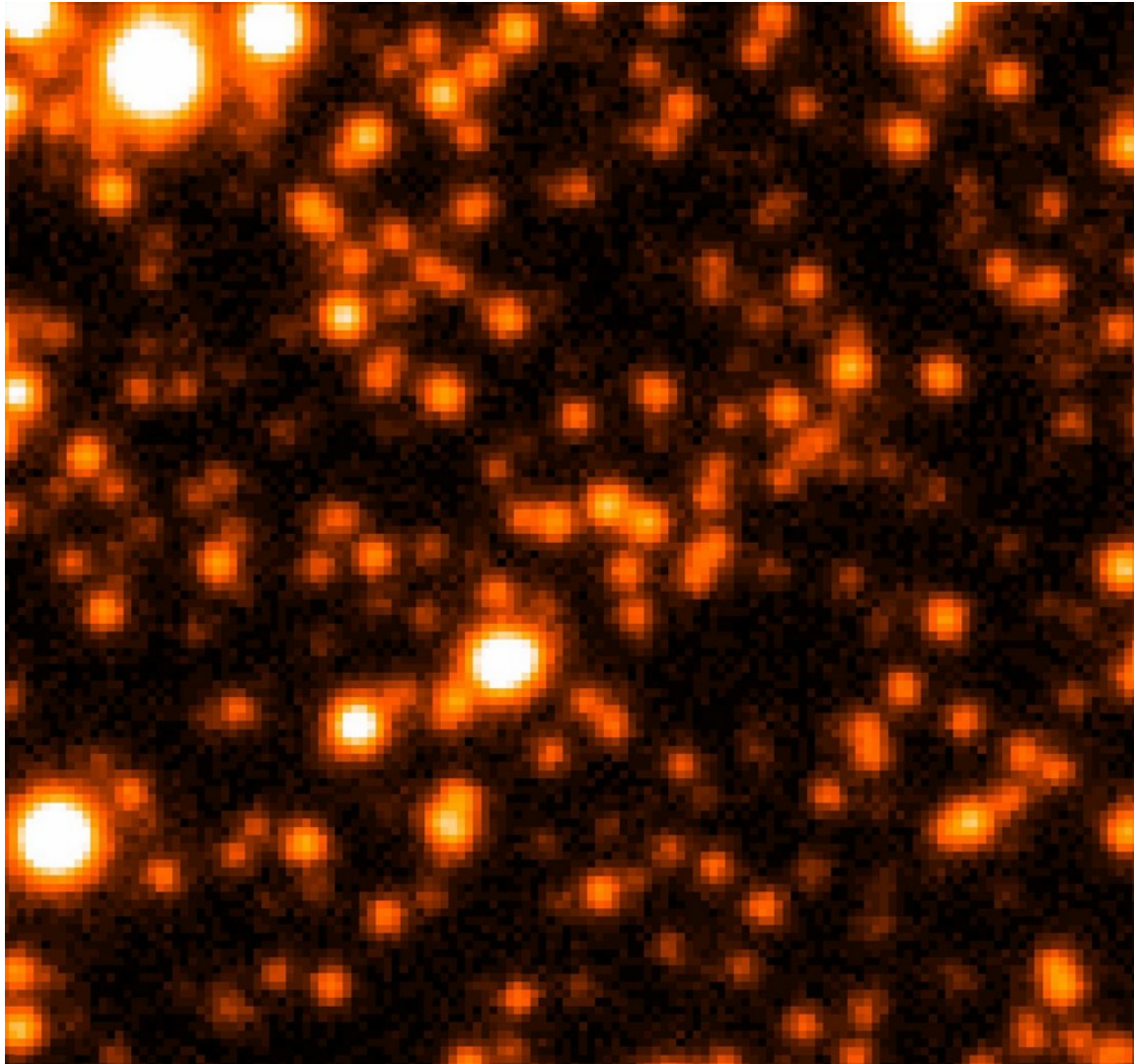


Globular Cluster NGC3201
Left: 4x4 FOVs BlueSi in
'MUSE' mode.

Right:
1 FOV real
MUSE data.

data+scene:
Sven Martens



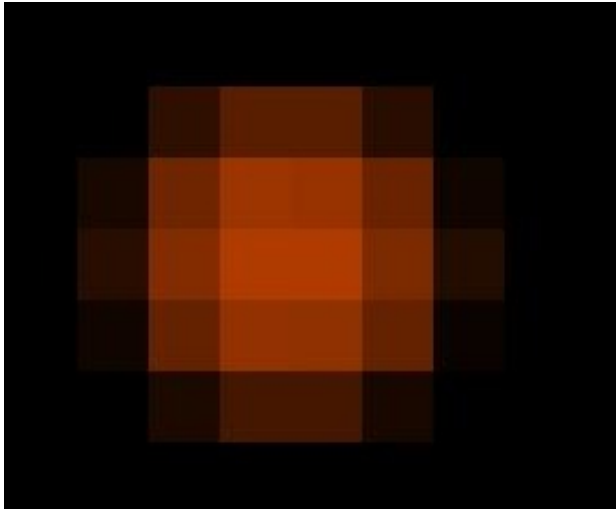


A single 'slice' with identical scale/cuts. Simulated vs real.

Why?

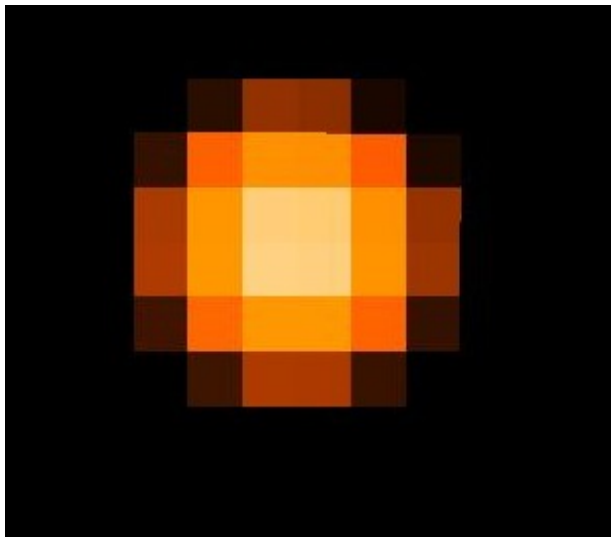
- science feasibility tests
- flexing tool chains and analysis tools
- tuning required S/N, sky, moon, Obs
- learning about BlueMUSE's strengths & quirks

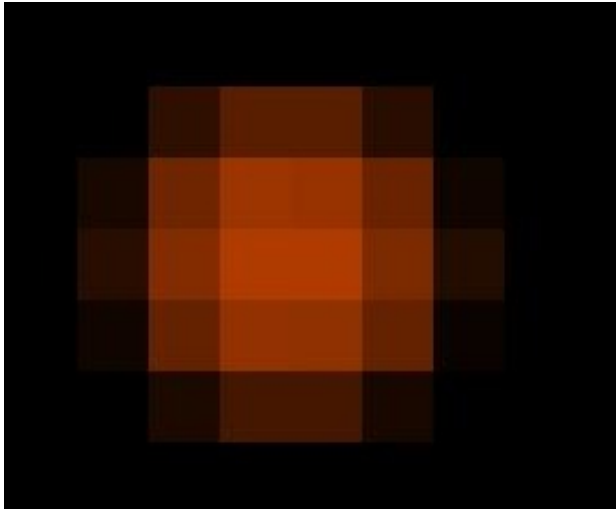




For example:
rectangular pixels.

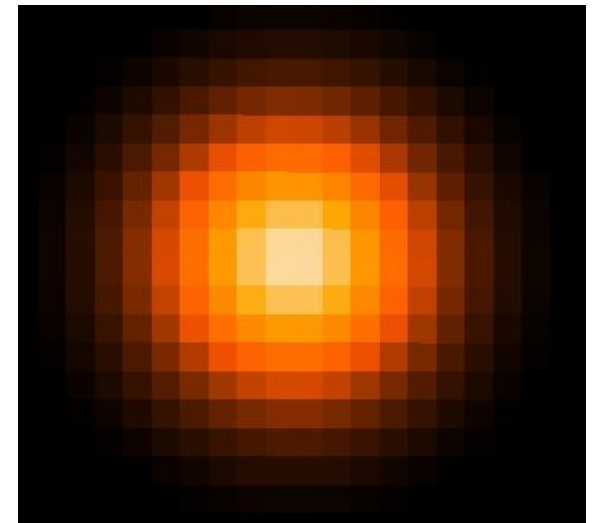
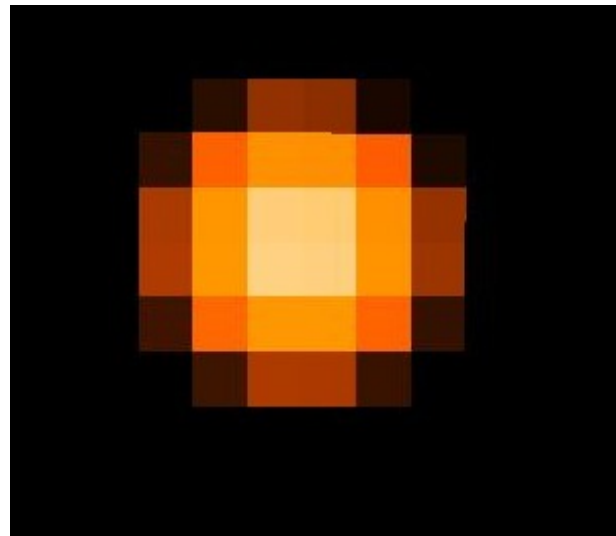
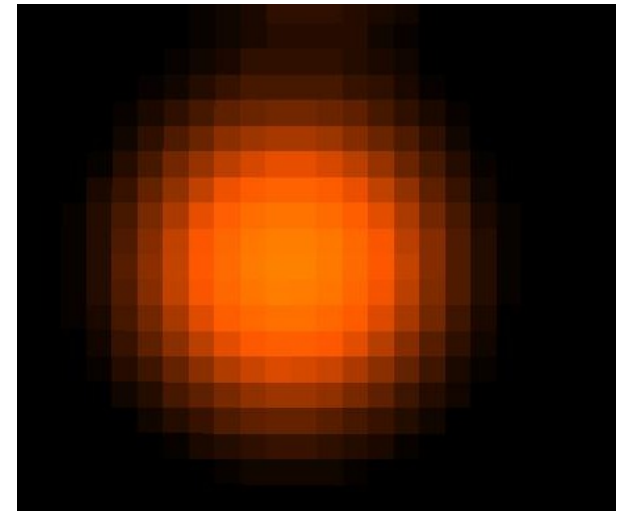
Top: single obs, 'in X'
Bottom: added 'X' and 'Y'

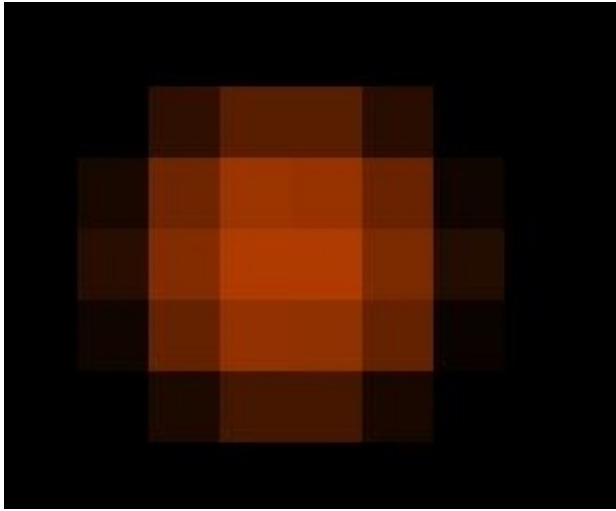




For example:
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Top: single obs, 'in X'
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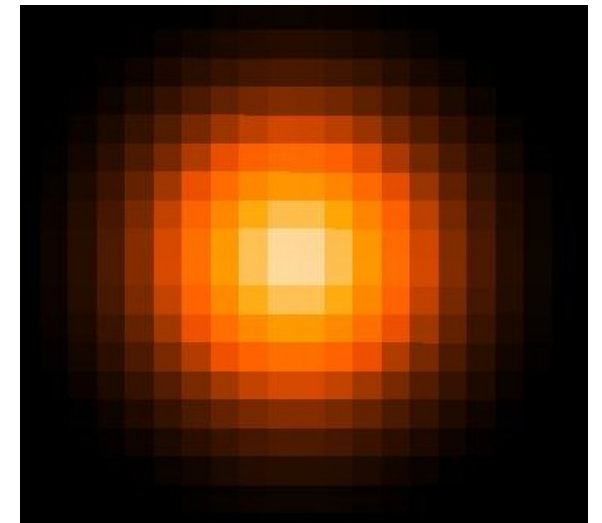
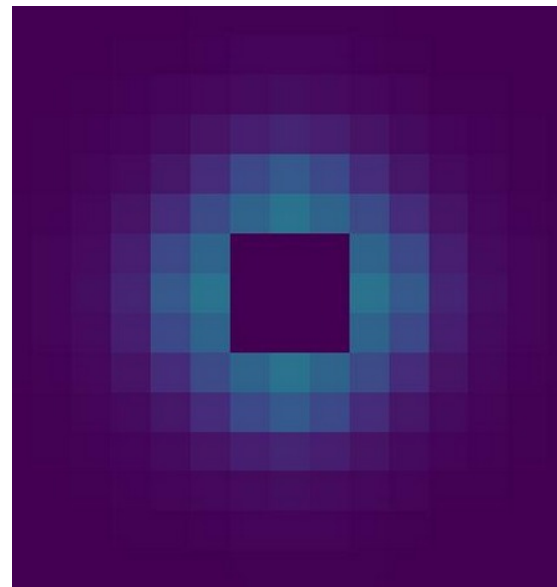
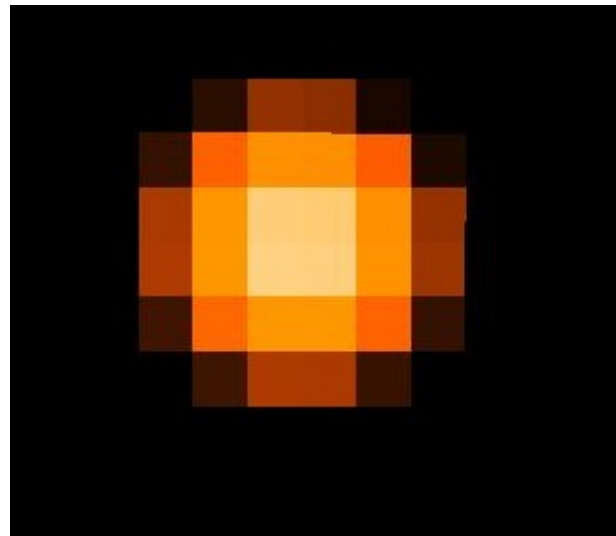
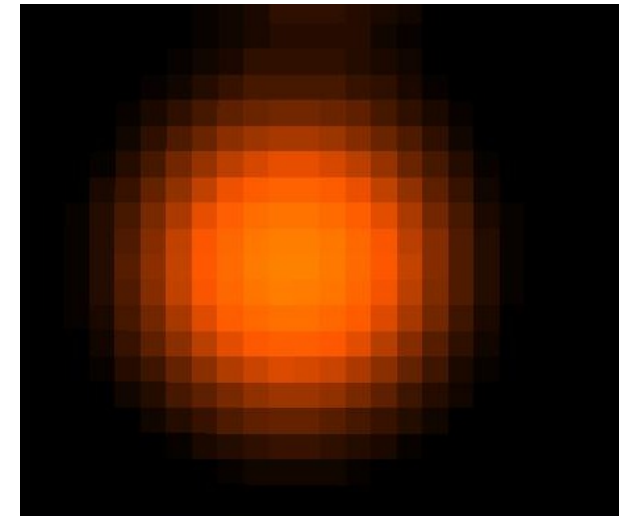




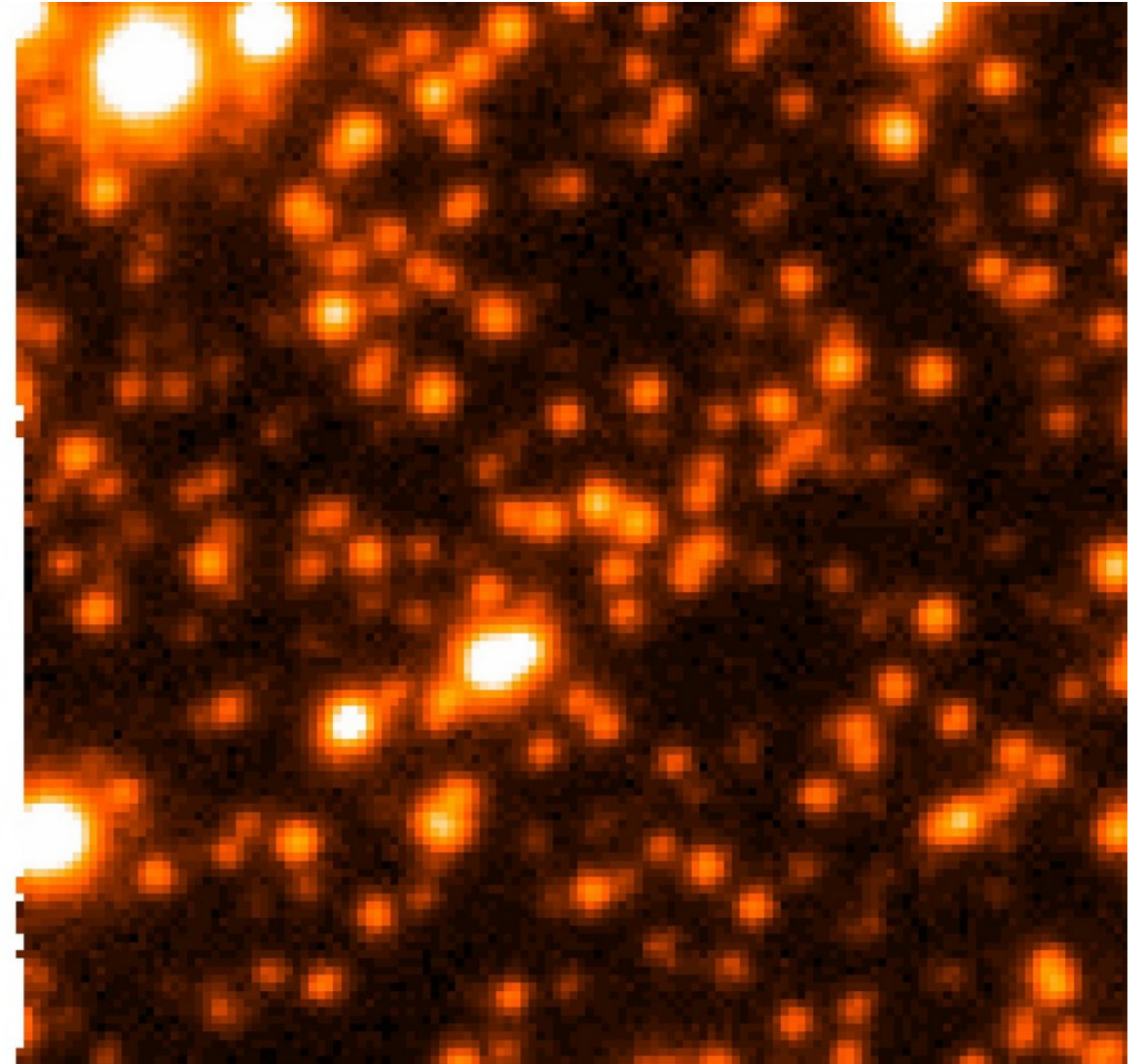
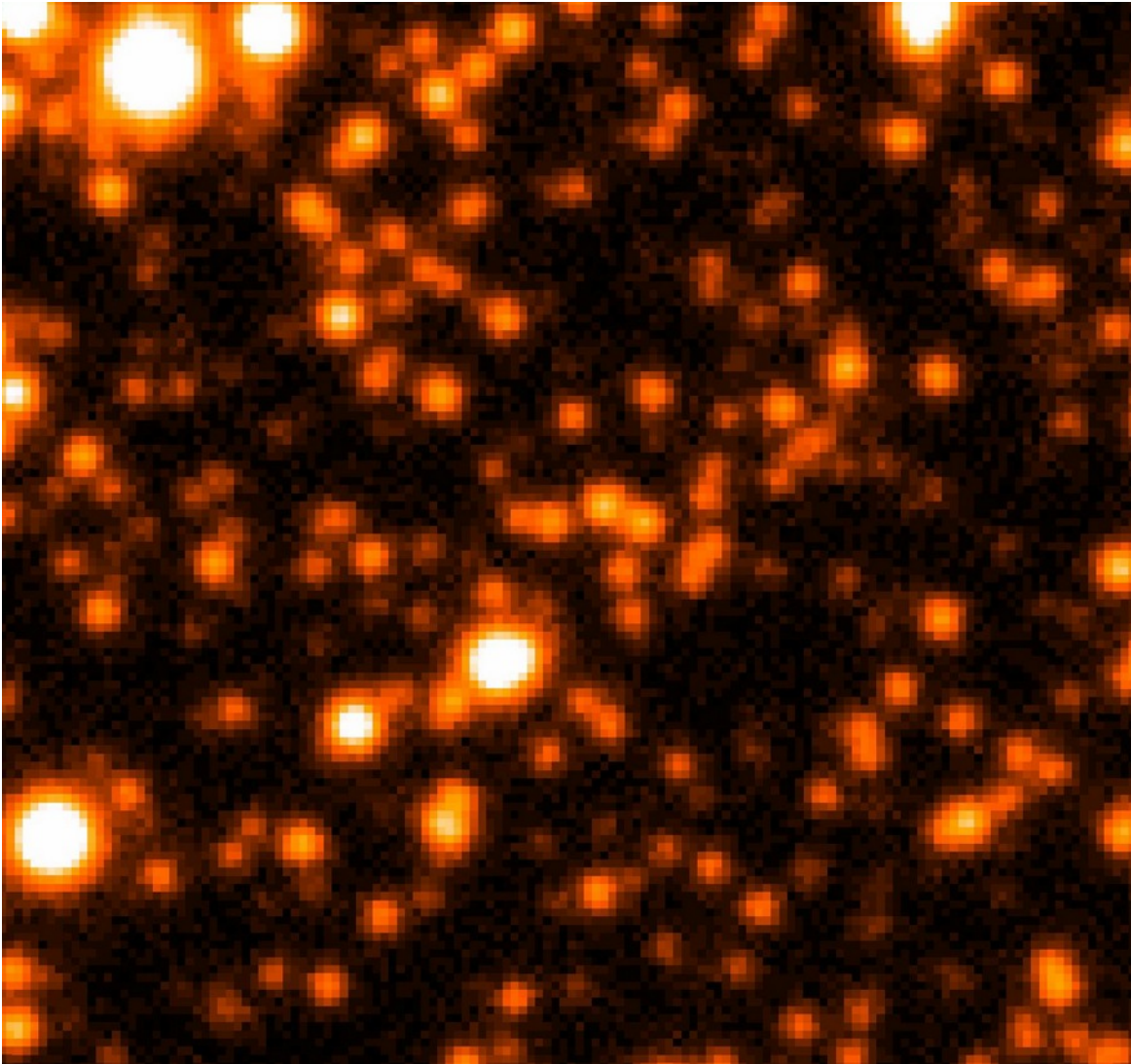
For example:
rectangular pixels.

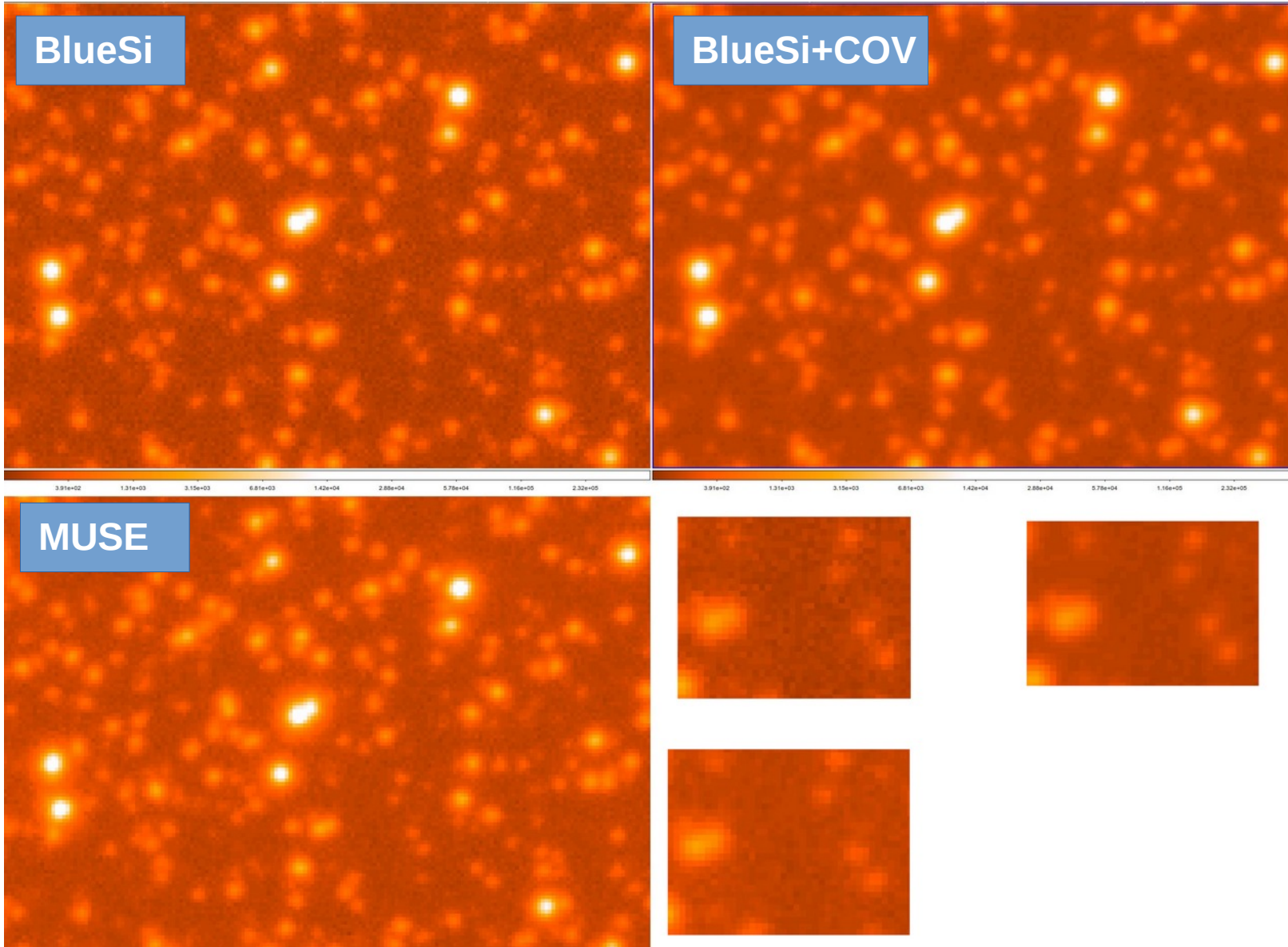
Top: single obs, 'in X'
Bottom: added 'X' and 'Y'

'X&Y' - square pixels



Example #2: noise characteristics.





example #2

Granularity of noise.
Impact of covariances
derived from 3D
autocorrelation in
observed noise.

Work in progress...
(+ DRS,
Peter Weilbacher)

Why?

- science feasibility tests
- flexing tool chains and analysis tools
- tuning required S/N, sky, moon, Obs
- learning about BlueMUSE's strengths & quirks

What (is being done)?

- a (Python) code reading in a scene definition and creating one (or several) data cubes
- ...coming up next...



What (is being done)?

```

1 wav_min = 3500.0           #lower wavelength limit (AA)
2 wav_max = 6000.0           #upper wavelength limit (AA)
3 R        = 4000.0» »      #Resolution at 5000 AA
4 d_lambda = 0.66           #constant bin size (AA)
5 pixsize  = 0.2            #detector pixel size (arcsec)
6 namebase = 'out/talk_small_blue' #name base of generated cubes
7 instrument = 'BlueMUSE'    #or BlueMUSE
8 seeing   = 0.9            #parameter for PSF (arsec)
9 airmass  = 1.1            #parameter for PSF
10 skytable = 'skytable_30dor.fits' #sky model
11 psf_beta = 2.5            #parameter for PSF
12 psf_l0   = 22            #parameter for PSF (m)
13 cpus     = 0             #number of cpus used
14 raw_dump = yes           #write cube before PSF/LSF/NOISE
15 var_ext  = yes           #add variance extension
16 do_lsf   = yes           #render and write LSF
17 debug_lsf = no          #plot LSF and insert delta peaks + fit
18 do_psf   = yes           #render and write PSF
19 do_noise = yes           #render and write final noise cube
20 add_sky  = no            #keep sky emission in data?
21 nexp     = 4             #number of exposures
22 texp     = 600           #exposures time per exposure (sec)
23 readout  = 3             #readout noise level (e-)
24 dark     = 3.0           #dark current (e-/hour)
25 saturation = 65535       #16bit saturation I assume
26 input_type = 'scene/scene_big30dor.conf' #scene/file or raw/file BlueSi cube
27 rectangular = 'yes'
28 rect_mode = 'x'

```

A simple external configuration file.
→ contains conditions

Usually not specific to the science case!



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```

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→ contains conditions

Usually not specific to the science case!

raw input or new scene?



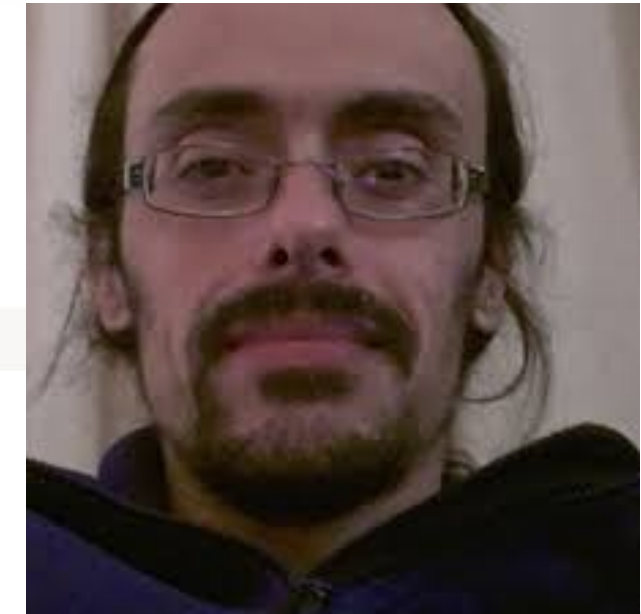
What (is being done)?

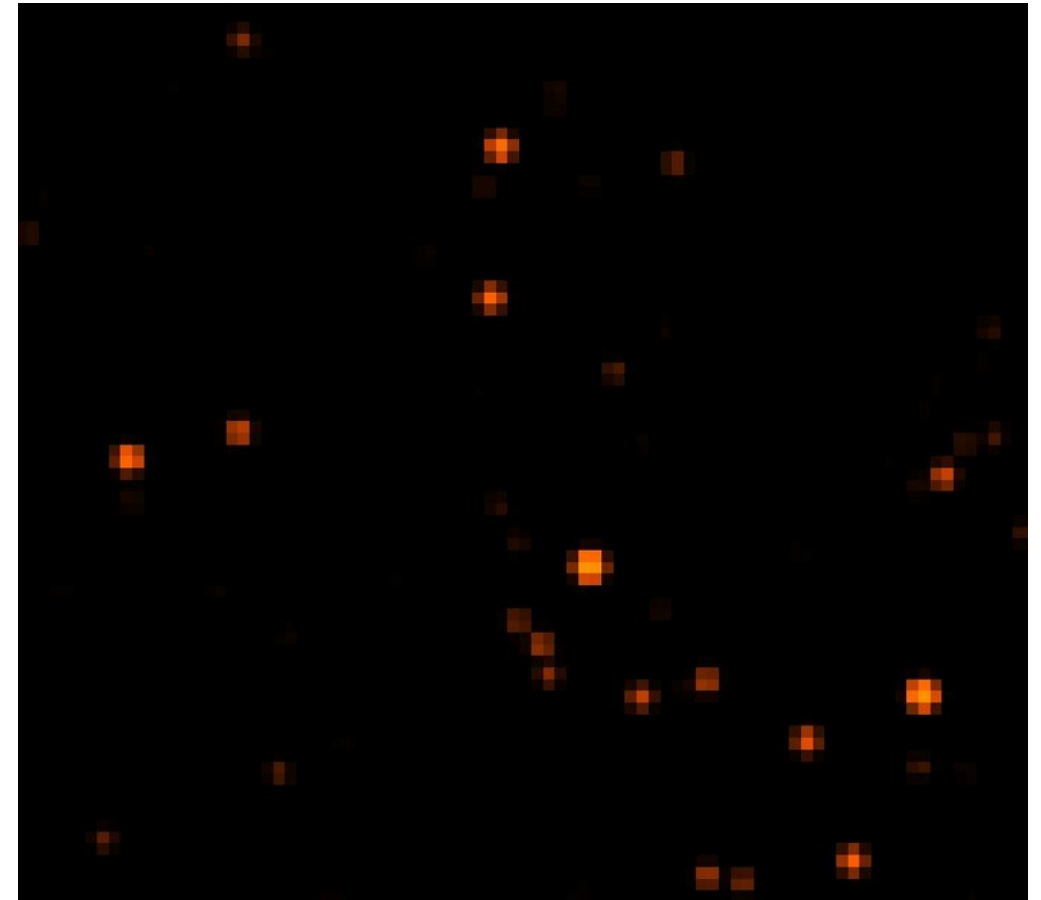
```

1 fov = 'all' # 'reg/file.reg' or 'all'
2 objectfile = '30dor/30Dor_BlueSi_HST_catV220224_compl.csv' #ASCII list
3 objectdir = '30dor/FITS' #directory of spectra
4 maxcount = 10 #only plot this number of objects, -1 = inf
5 vrad = yes #apply radial shifts from table
6 vac2air = no #apply air refraction to model spectra
7 extinction = yes #apply extinction to whole frame
8 Ebv = 0.39 #E(B-V)
9 simulated_object = '30Dor' ←
10 model_spectra_packaging = 'Rodriguez,Norberto'
11 scene_file_creator = 'Wendt,Martin'
12 scene_type = '30dor'
13

```

Another file: describing the science scene.
(unless a 'raw' cube is being used)





A 'raw' cube:
no LSF, PSF, sky, noise
(sub-pixel placement)

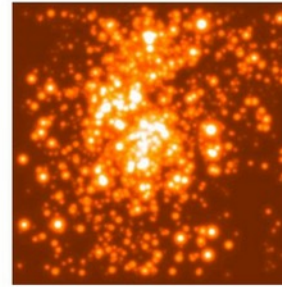
1 Mpc



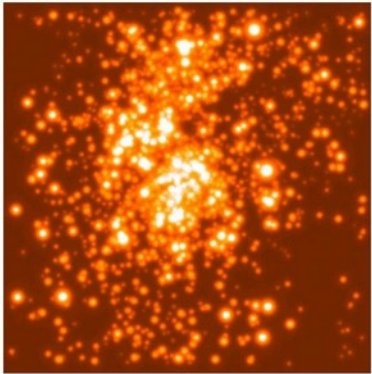
0.5 Mpc



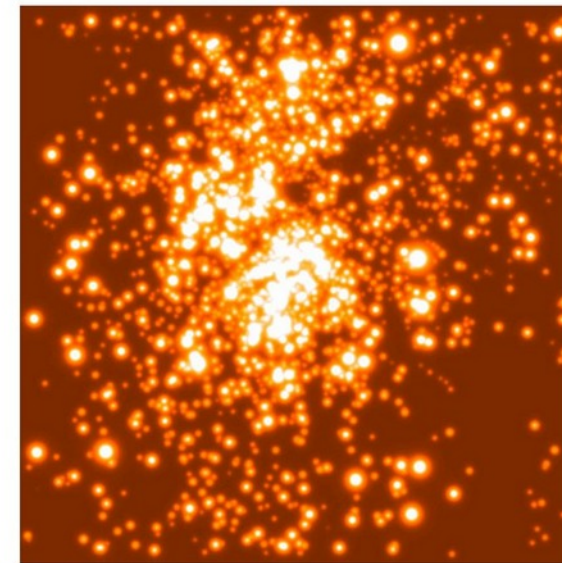
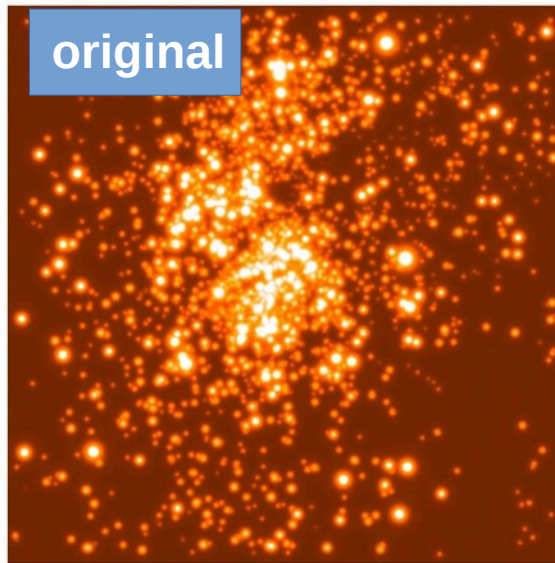
0.1 Mpc



0.075 Mpc



original



BlueSi:
full cubes of
30 Dor at
different
distances
In MUSE mode.

Bottom right:
BlueMUSE mode

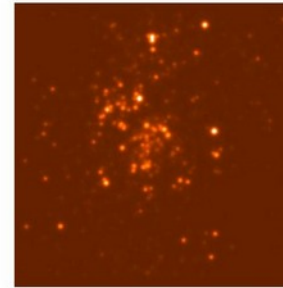
1 Mpc



0.5 Mpc

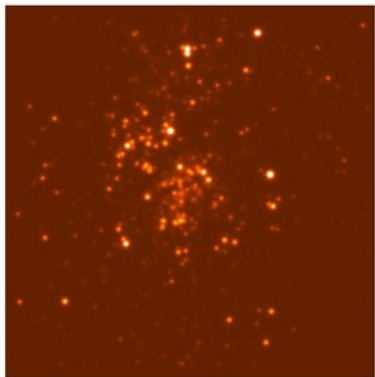


0.1 Mpc

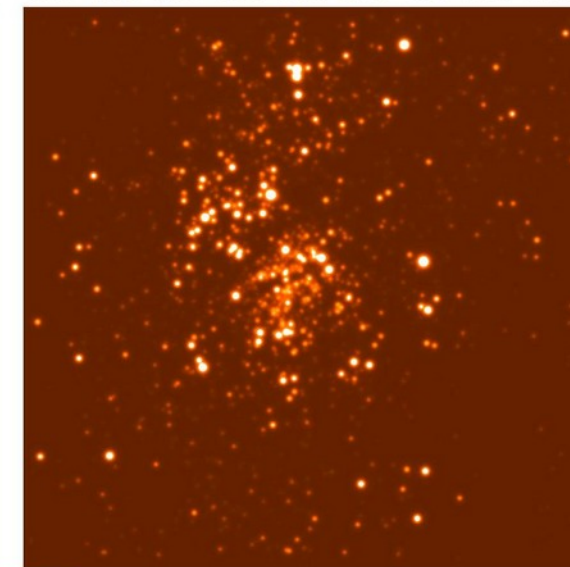
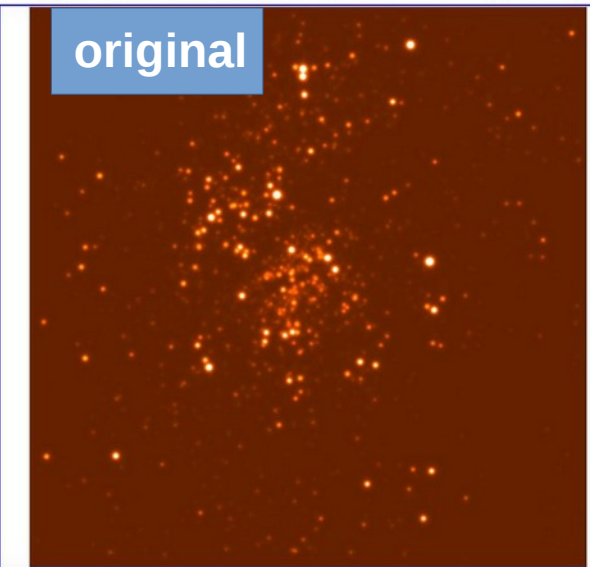


here in linear scale.

0.075 Mpc



original



(magnitudes scaled with distance by Norberto)

Why?

- science feasibility tests
- flexing tool chains and analysis tools
- tuning required S/N, sky, moon, Obs
- learning about BlueMUSE's strengths & quirks

What (is being done)?

- a (Python) code reading in a scene definition and creating one (or several) data cubes
- ...coming up next...

How can my science interest be covered?



What (is being done)?

```
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```



Another file: describing the science scene.
(unless a 'raw' cube is being used)

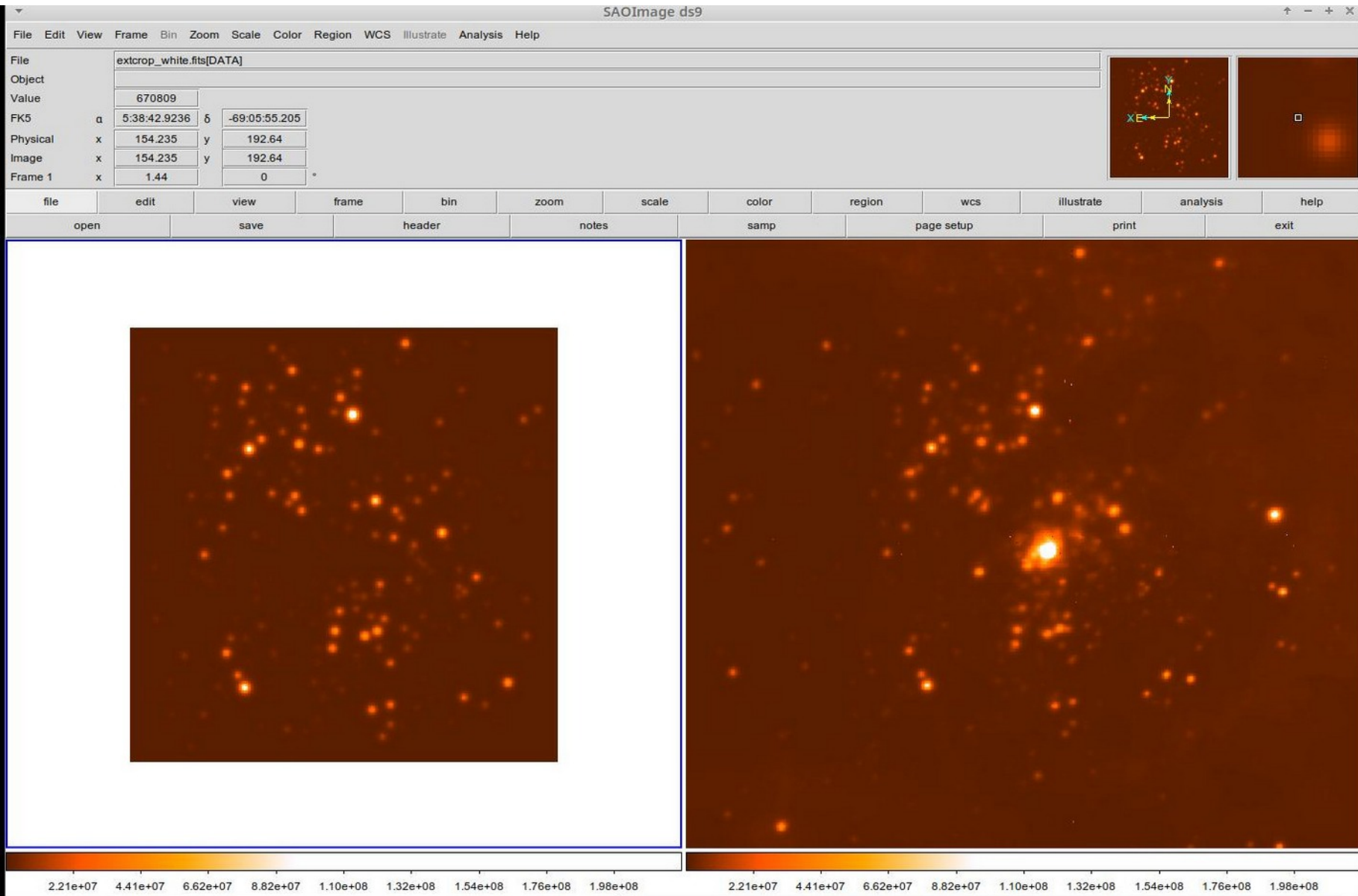
```
1 #HTTP,RAJ2000,DEJ2000,F555mag,vsini,vra,model
2 053853.129-690702.50,84.721371,-69.117361,18.42099952697754,10.0,265.0,BG19000g425v2.vis.7.fits
3 053853.078-690703.49,84.721158,-69.117636,20.761999130249023,160.0,275.0,lte09400-4.50-0.0.PHOENIX-ACES-AGSS-COND-2011-HiRes.fits
4 053850.694-690703.33,84.711225,-69.117592,17.864999771118164,150.0,265.0,BG21000g425v2.vis.7.fits
5 053850.018-690705.32,84.708408,-69.118144,17.216999053955078,10.0,265.0,BG24000g425v2.vis.7.fits
6 053849.927-690705.28,84.708029,-69.118133,17.13599967956543,280.0,305.0,BG25000g425v2.vis.7.fits
7 053849.660-690704.32,84.706917,-69.117867,20.481000900268555,70.0,265.0,lte10000-4.50-0.0.PHOENIX-ACES-AGSS-COND-2011-HiRes.fits
8 053848.435-690704.69,84.701812,-69.117969,18.312999725341797,180.0,235.0,BG20000g425v2.vis.7.fits
9 053852.699-690651.55,84.719579,-69.114319,22.80500030517578,130.0,295.0,lte06900-4.50-0.0.PHOENIX-ACES-AGSS-COND-2011-HiRes.fits
10 053847.602-690705.85,84.698342,-69.118292,17.128999710083008,420.0,285.0,BG25000g425v2.vis.7.fits
11 053849.853-690657.94,84.707721,-69.116094,18.555999755859375,10.0,295.0,BG19000g425v2.vis.7.fits
```

A plain CSV input file + all the required synthetic spectra

- wavelength coverage BlueMUSE (+MUSE)
- spectral resolution / sampling
- air / vacuum

+ physical models / algorithms

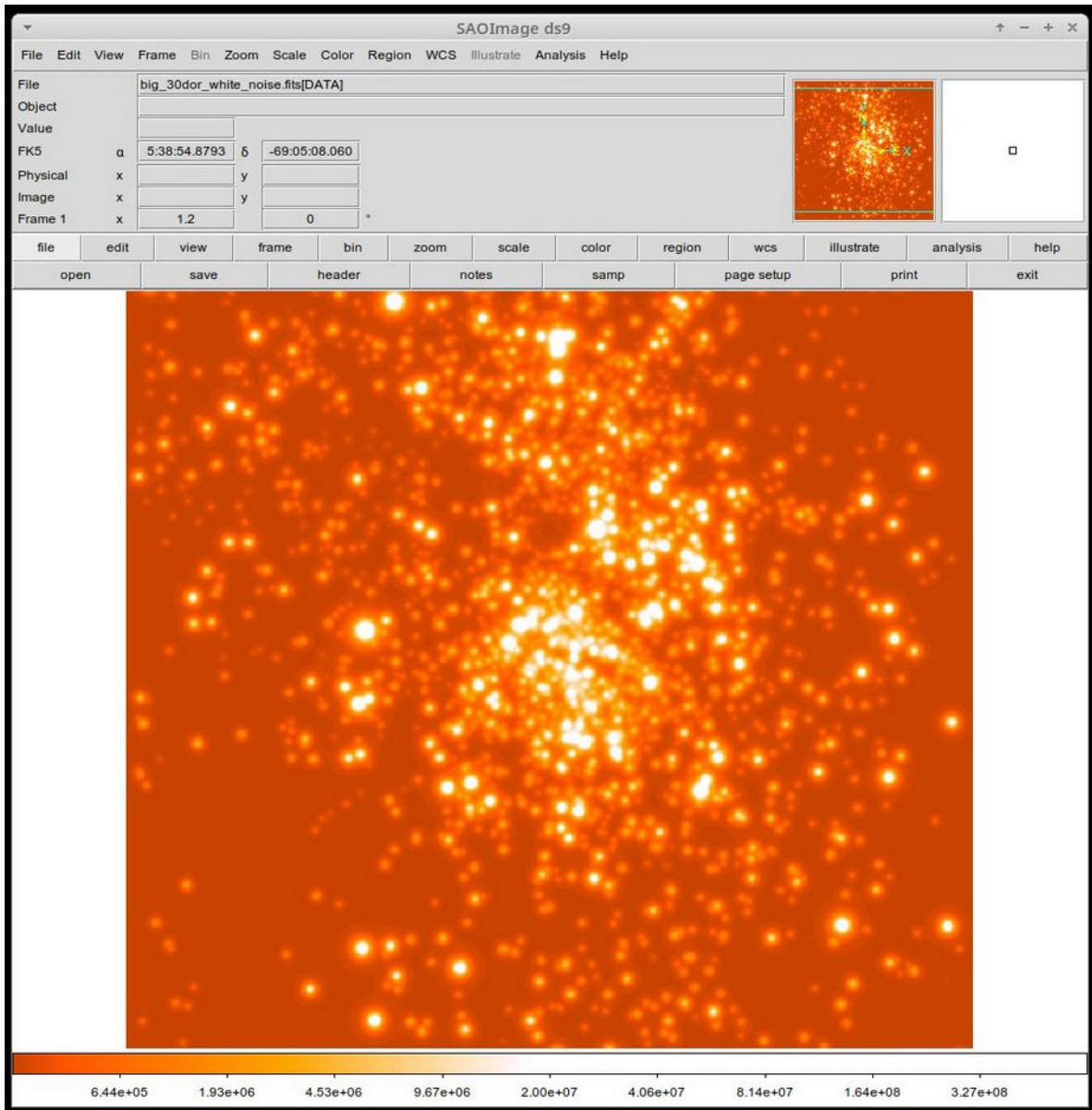
- extinction (universal)
- rotation broadening?
- LAE model? (Tanya, John)



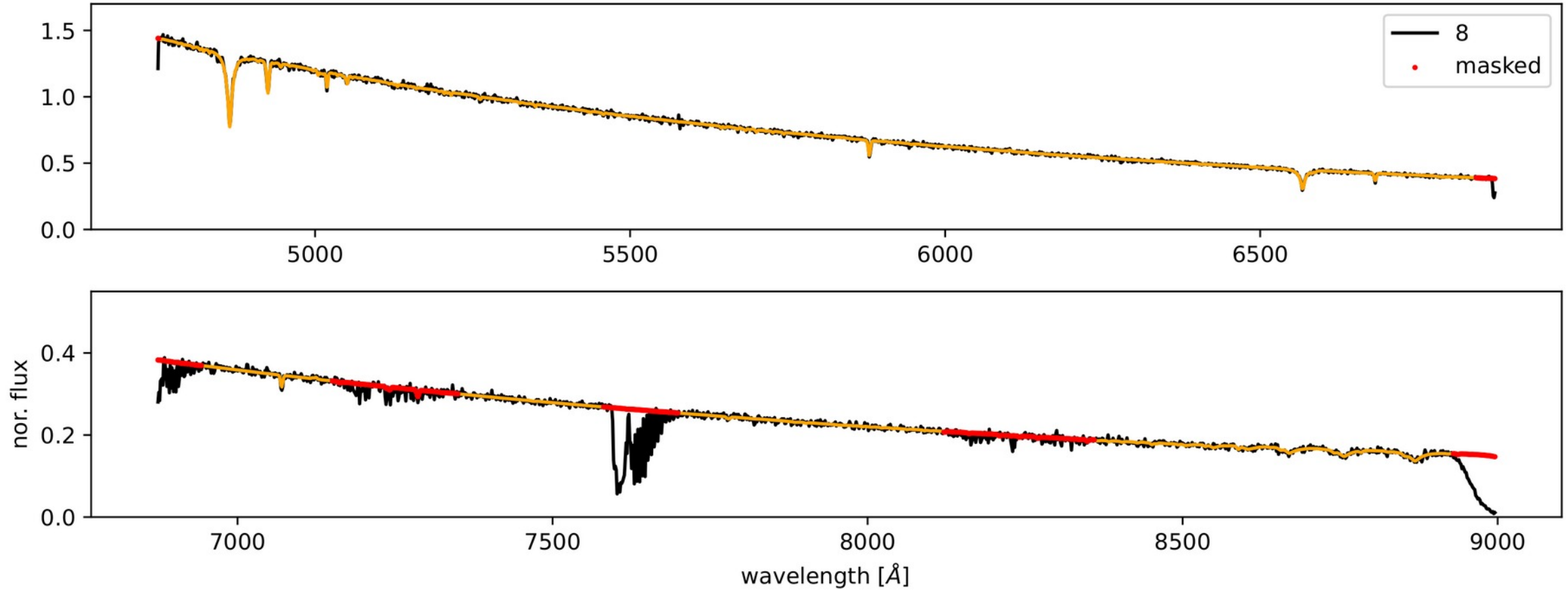
30 Dor scene #1
from Norberto

left: BlueSi
right: MUSE

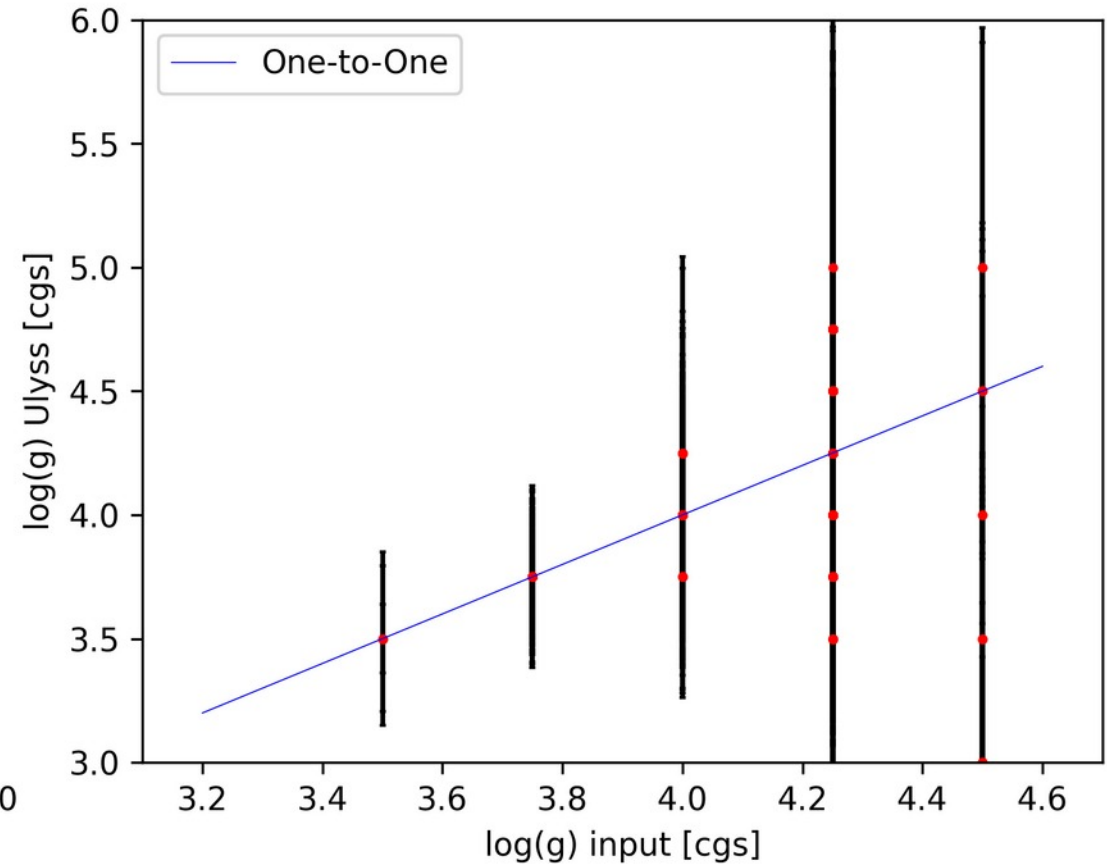
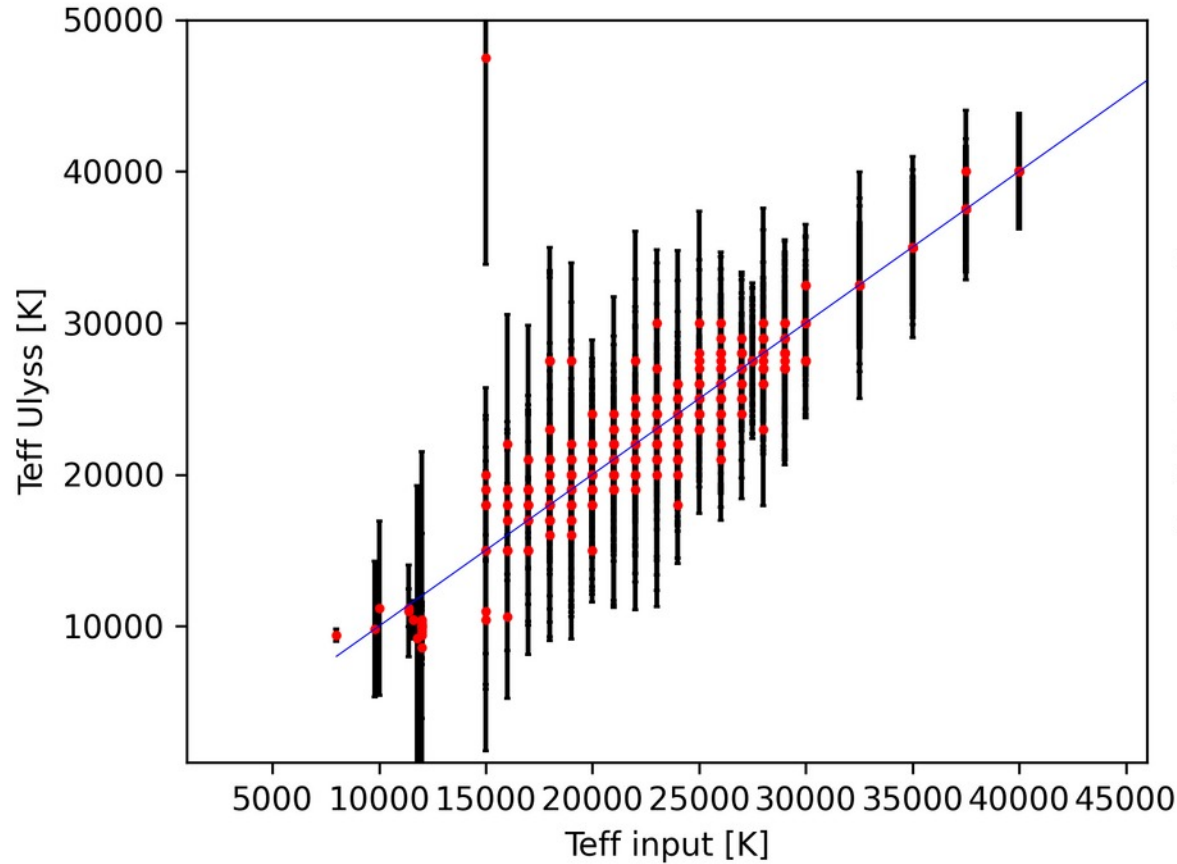
Scene based on
fully analysed
objects only.



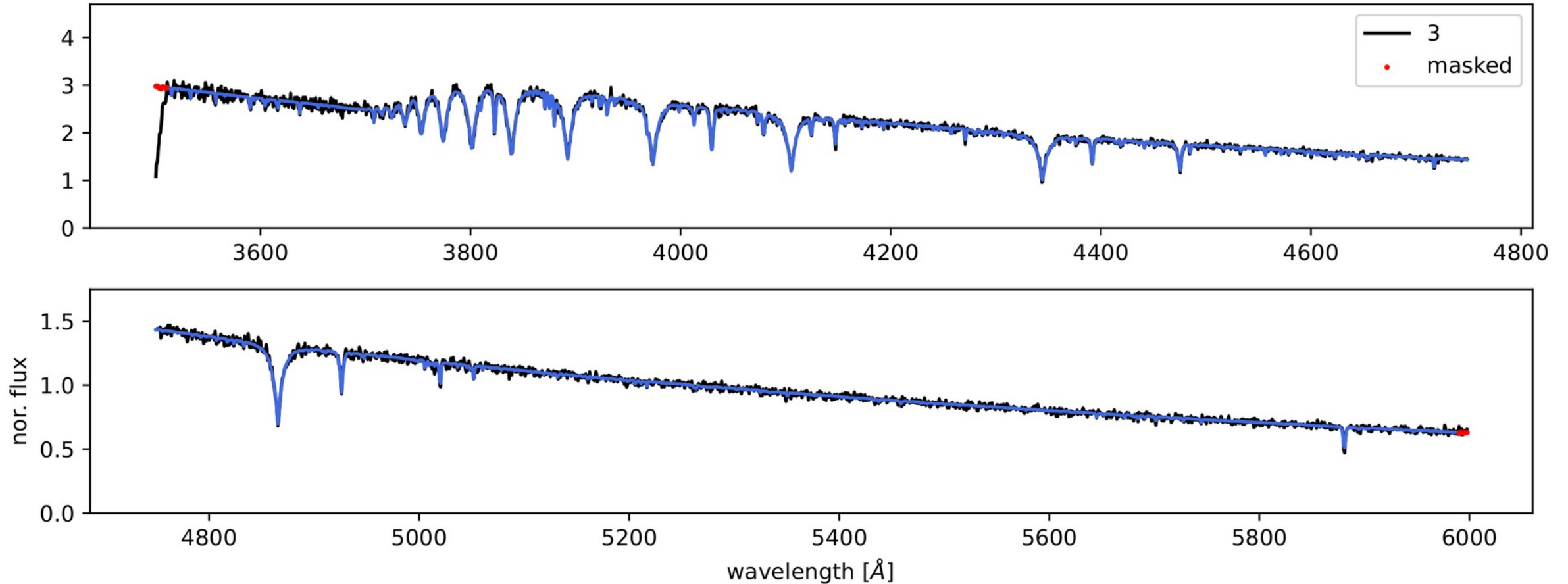
BlueSi now:
 based on photometric HST
 catalog.
 Estimation of the stellar
 parameters based on the colors
 and synthetic isochrones
 (no T_{eff} , $\log(g)$).
 ~ 3,000 stars (Tlusty + PHOENIX)



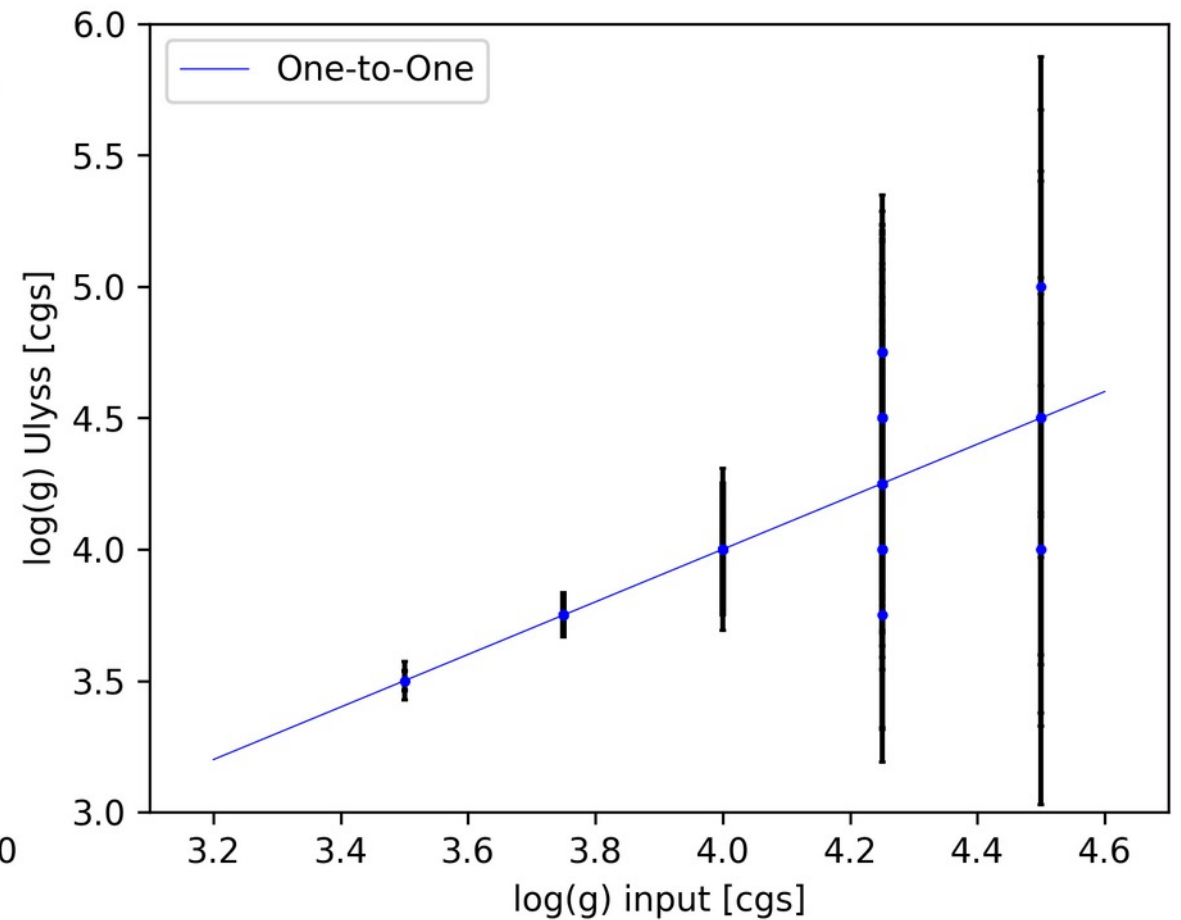
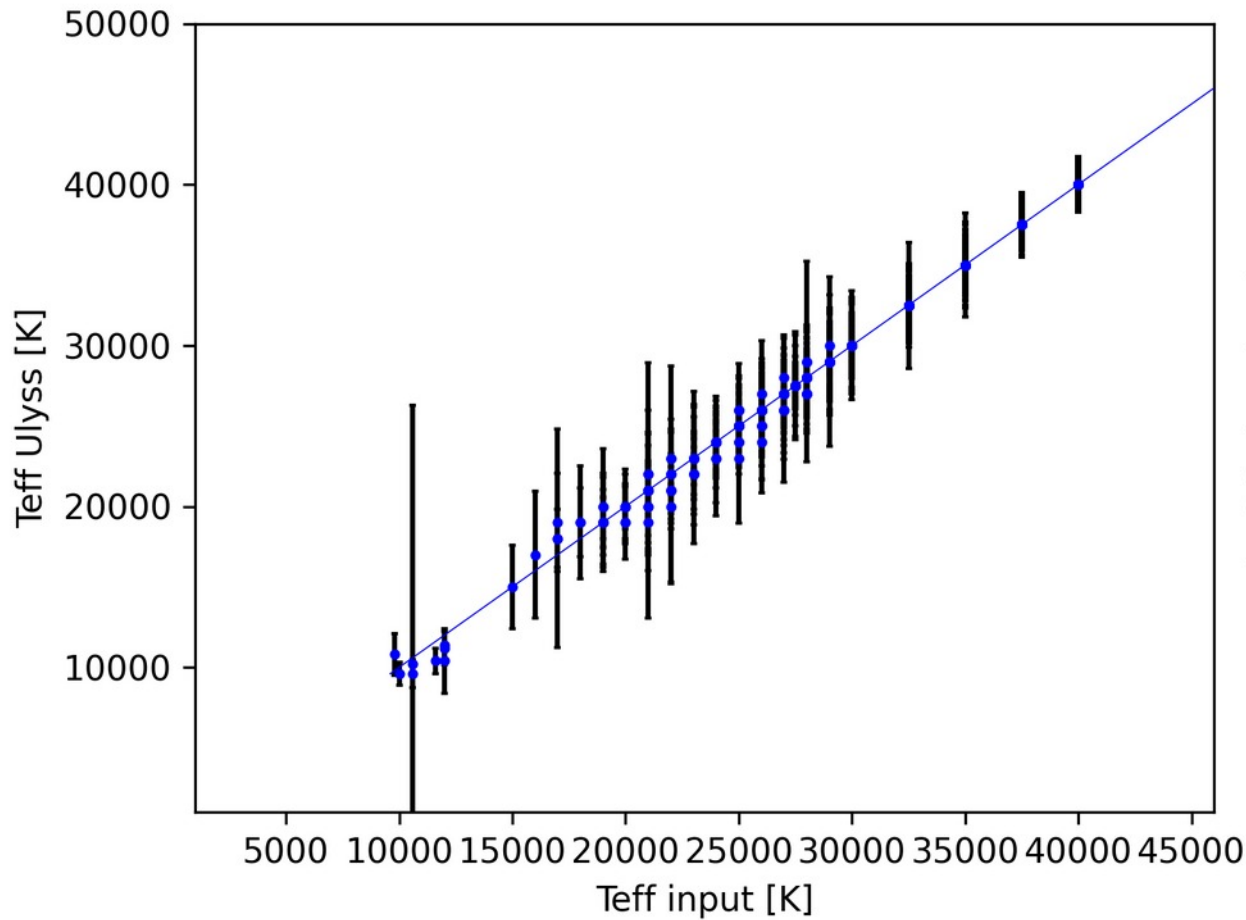
Exemplary extracted spectrum from BlueSi cube in MUSE mode and model fit by Norberto.



First input-output comparison for MUSE specs by Norberto.



Exemplary extracted spectrum from BlueSi cube in amazing BlueMUSE mode and model fit by Norberto.



Input-output comparison for the BlueMUSE cube by Norberto.

What is next?

- test methods/tools/science on existing simulations (stars)
- extend 30 Dor (diffuse gas)
- CREATE NEW SCIENCE objects (SpinSim?, other input)
- Simulation of:
 - specific technical parameters
 - * ADC?
 - * spectral/spatial variation LSF/PSF
 - * rectangular pixels
 - specific observing conditions
 - * exposure times
 - * sky models

